

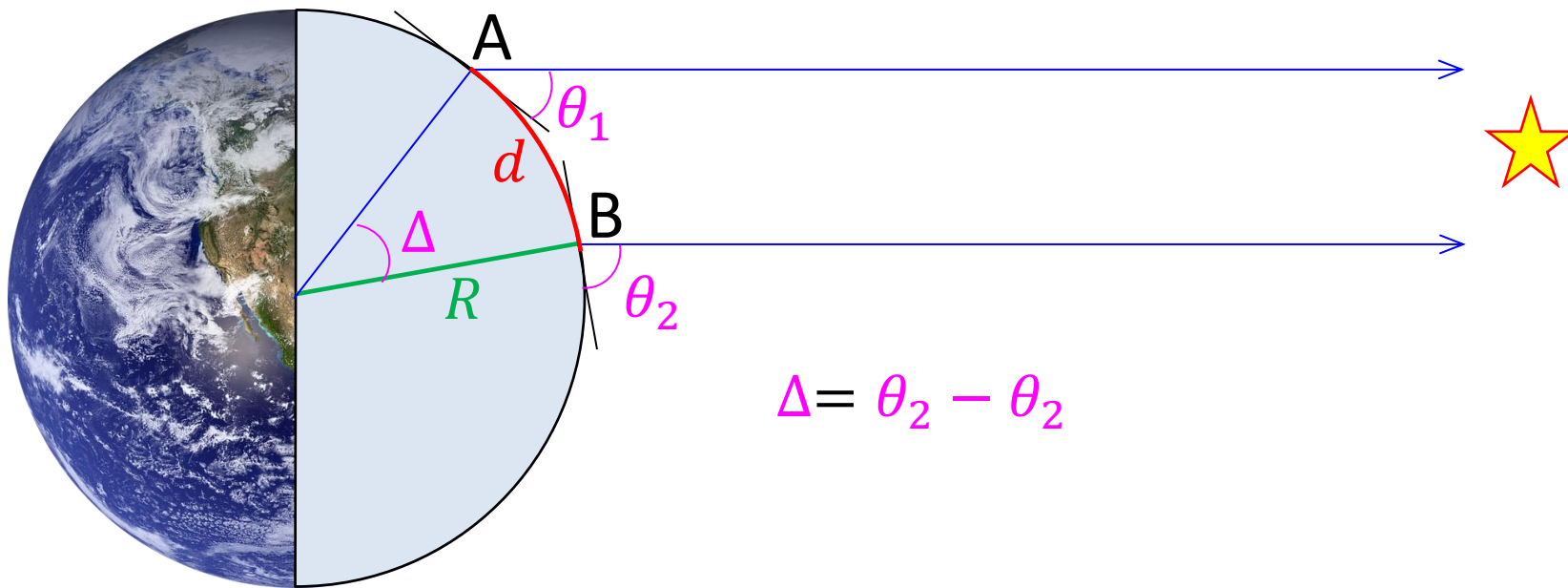
地球の半径を求める

方法1

$$R = \frac{\text{地球の赤道の長さ}}{2\pi} = \frac{40,000\text{km}}{2\pi} \doteq 6,370\text{km}$$

方法2

$$R = \frac{d}{\Delta}$$



方法1

地球の質量を求める

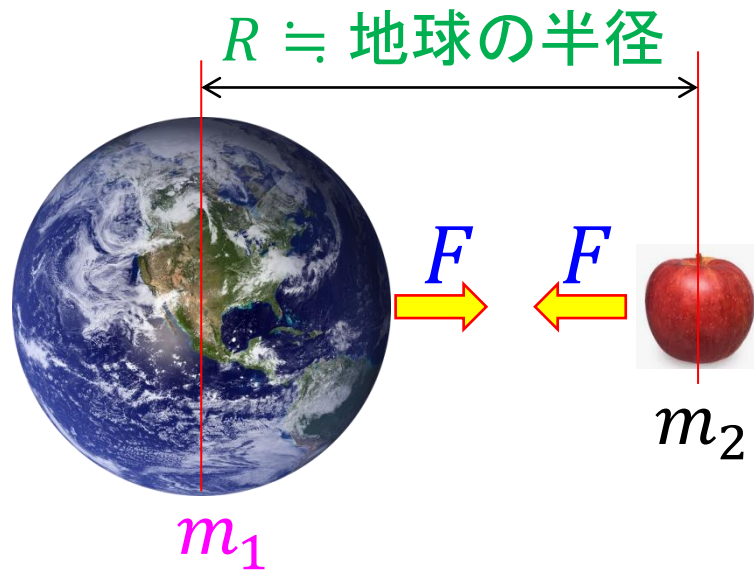
$$F = G \frac{m_1 m_2}{R^2}$$

$$F = m_2 g$$

$$G \frac{m_1 m_2}{R^2} = m_2 g$$

$$\begin{aligned} m_1 &= \frac{g R^2}{G} \\ &= \frac{9.807}{6.67 \times 10^{-11}} \times (6378 \times 1000)^2 \\ &\doteq 6.0 \times 10^{24} \text{ kg} \end{aligned}$$

万有引力定数 $G: 6.67 \times 10^{-11} \text{ m}^3 \cdot \text{kg}^{-1} \cdot \text{s}^{-2}$
重力加速度 $g: 9.807 \text{ m} \cdot \text{s}^{-2}$



方法2

地球の質量を求める

$$\text{万有引力 } F = G \frac{m_1 m_2}{R^2}$$

$$\text{遠心力 } F = m_2 r \omega^2 = m_2 r \left(\frac{v}{R} \right)^2 = \frac{m_2 v^2}{R}$$

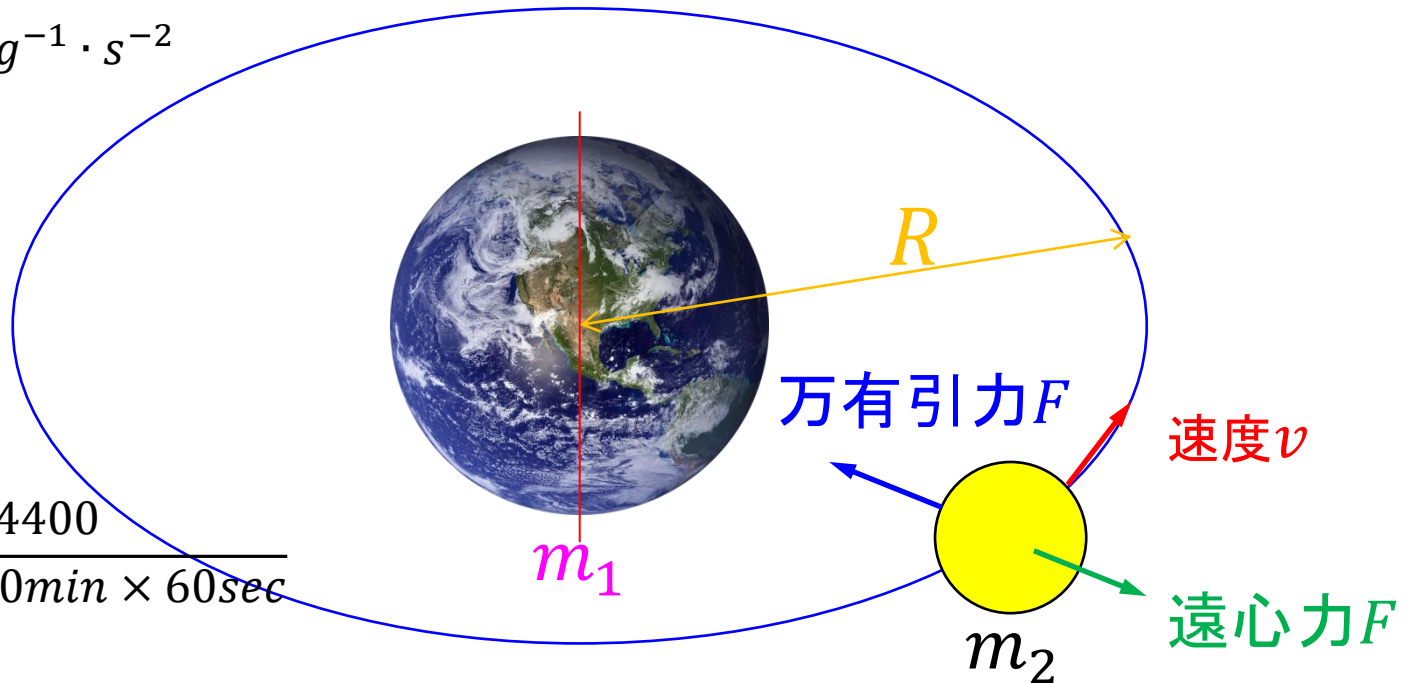
$$\text{遠心力 } F = \text{万有引力 } F$$

$$\frac{\cancel{m_2} v^2}{R} = G \frac{\cancel{m_1} \cancel{m_2}}{R^2}$$

$$m_1 = \frac{R v^2}{G} = \frac{384400 \times 10^3 \times (1.02 \times 10^3)^2}{6.67 \times 10^{-11}} \doteq 6 \times 10^{24} \text{ kg} = 6 \times 10^{21} \text{ トン}$$

$$G = 6.67 \times 10^{-11} \text{ m}^3 \cdot \text{kg}^{-1} \cdot \text{s}^{-2}$$

$$R = 384,400 \text{ km}$$



$$v = \frac{2\pi \times 384400}{27.3 \text{ 日} \times 24 \text{ h} \times 60 \text{ min} \times 60 \text{ sec}} \doteq 1.02 \text{ km/sec}$$

太陽の質量を求める

$$m_1 = \frac{Rv^2}{G} = \frac{1.5 \times 10^8 km \times 10^3 \times (30 \times 10^3)^2}{6.67 \times 10^{-11}} \doteq 2.0 \times 10^{30} kg$$

$$G = 6.67 \times 10^{-11} m^3 \cdot kg^{-1} \cdot s^{-2}$$

$$R = 1.5 \times 10^8 km$$

$$v = \frac{2\pi \times 1.5 \times 10^8 km}{365 \text{ 日} \times 24h \times 60min \times 60sec} \doteq 30.0 km/sec$$

	太陽	地球	月
質量 kg	2.0×10^{30}	6×10^{24}	7.35×10^{22}
体積 km^3	1.43×10^{18}	1.08×10^{12}	2.40×10^{10}
密度 g/cm^3	1.4	5.5	3.1